

# How to Work with the Programmable Spectrum and Example (Black-Body Radiation)



Providing maximum versatility for your optical simulations is one of our most fundamental objectives. In this document we show you how to work with the Programmable Spectrum: that is, how to define a function that assigns a different complex weight to each wavelength/frequency present in the spectral make-up of a field, working under assumptions of stationary behaviour. The black-body emitter is one of the default spectrum models in VirtualLab, but we use it here as a basic programming example.

## Where to Find the Programmable Spectrum

dit Programmable Spectrum Generator X	1 4 1 4 1 4 1 4 1 4 1 4 A	29
Generator Settings <ul> <li>Specification in Wavelength Domain</li> <li>Central Wavelength</li> <li>Stope</li> </ul>	Spectral Spectral Spectral Spectral Spectral	provide Ge
Definition	Source Code Editor	1 • ×
Numerical Settings	opping       1       2       Complex value = new Complex(0,0);       Positio         1       2       3       ************************************	ın [double]
Size of Wavelength Window 200 nm Number Data Points 50		

# **Setting Up the Sampling**



# Writing the Code



- The panel on the right shows a list of available independent parameters.
- Position represents the independent variable (either wavelength or frequency, as pre-set in the configuration dialogue).
- The code in the Main Function must return a Complex value per Position, which is determined by the function programmed by the user.
- Use the Snippet Body to group parts of the code in support functions.
- The final sampling of the function is determined by the settings from the previous dialogue.

#### Output

- The output is a one-dimensional graph of the programmed complex-valued function.
- It is possible to use the generated spectrum as the spectral make-up of the source in your Optical Setup.
- The number of separate spectral modes when the programmed spectrum is used in a source is equivalent to the number of samples in the spectrum.



#### **Programming a Black-Body Spectrum**

The power density associated to each wavelength (spectral density) when an emitter is assumed to behave like a black body at a certain temperature T is given by Planck's Law:

$$S(\lambda) = \frac{8\pi hc}{\lambda^5} \frac{1}{\exp\left(\frac{hc}{\lambda kT}\right) - 1} \tag{1}$$

- $S(\lambda) \rightarrow$  Spectral density
- $\lambda \rightarrow \text{Wavelength}$
- $h \rightarrow \ {\rm Planck's \ constant}$
- $c \rightarrow \ {\rm Speed}$  of light in vacuum
- $k \rightarrow \text{ Boltzmann constant}$
- $T \rightarrow$  Absolute temperature of black-body emitter

#### **Black-Body Radiation**

The maximum of the curve is achieved for the wavelength

$$\lambda_{\max} = \frac{b}{T} \tag{2}$$

where  $b = 2.8977729 \times 10^{-3} \,\mathrm{m\,K}$  represents Wien's displacement constant.

# **Programmable Spectrum: Setting Up the Sampling**

Edit Programmable Spectrum Generator		X III 81: Black Body Power Spectrum
Generator Settings		Diagram Table Value at x-Coordinate
Specification in Wavelength Domain     Central Wavelength	O Specification In Frequency Domain	
Shape Definition		6 0.7 0.8
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Number of Data Points = 2	00	0.15 0.2
Numerical Settings	Number Data Points 200 🖨	
O Sampling Distance 5 nm		to that the N



## **Programmable Spectrum: Entering the Programming Interface**

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Generator Settings <ul> <li>Specification in Wavelength Domain</li> <li>Central Wavelength</li> </ul>	O Sr	Source Code E	ditor Global Parameters Snippet Help Advanced Settin	ıgs	- D	×
Shape Definition		Snippet Body Main Function 9 5 7 1	<pre>Complex value = new Complex(0,6 /************************************</pre>	<pre>3); ************************************</pre>	Position [double]	
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Size of Wavelength Window     Sampling Distance	1 µm Numb 5 nm	er D	Check Consistency Validity: 🚹 🚺	ОК	Cancel	Help

## **Programmable Spectrum: Global Parameters**

- Once you have triggered open the Edit dialogue, go to the Global Parameters tab.
- There, Add and Edit two global parameters:
  - double TemperatureKelvin (0, NaN): represents the absolute temperature at which the black body is radiating.
  - bool Normalize: will the function be scaled so that the maximum allowed amplitude value is 1 (true) or not (false)?

Source Code	Editor					-		×
Source Code	Global Parameters	Snippet H	elp Advanc	ed Settings	3			
General Par	rameters							
Variable N	lame		Гуре			Description		
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# **Programmable Spectrum: Snippet Help**

Source Code	Editor						_		×
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Author					Last Modified	14/08/20	18		
spectrum a	coording to the black hould be scaled so th	body curve. Th	eureer can input whi weight is unity.	ch temperature ti	he black body sh	nall radiate a	at, and wi	hether th	le
Preview									
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Versi Last I	on: 1.0 Modified: Tues	sday, Augu	st 14, 2018						١.
Black	body radiation	is one of the snipr	he best-knowr	theories in	physics, ar	nd it is e to the bl	mploy ack-bo	ed	~
1	Check Consistency	Validity: 🕑			ОК	Ca	ncel	H	elp

- **Optional:** you can use the Snippet Help tab to write instructions, clarifications, and some metadata associated to your snippet.
- This option is very helpful to keep track of your progress with a programmable element.
- It is especially useful when the programmable element is later disseminated to be handled by other users!

Hint: Use HTML commands to format the text

# **Programmable Spectrum: Snippet Help**

Source Code Editor	— <b>—</b>	×	Spinnet Help		пх
Source Code Global Parameters Snippet Help /	Advanced Settings		Shipper neip		
Title Black-Body Spectrum	Version 1.0		Black-Body S	pectrum	~
Author	Edit Programmable Spectrum Generator	×	Diack Douy o	poorum	
Black-body radiation is one of the best-known th spectrum according to the black-body curve. Th spectrum should be scaled so that the maximum	Generator Settings	ecification In Frequency Domain	Version: 1.0 Last Modified: Tueso Black-body radiation in employed across seven the black-body curve.	day, August 14, 2018 s one of the best-known theories in physics, and it is eral fields. This snippet generates a spectrum accordin The user can input which temperature the black body s	ng to shall
	Perintion Zeriation	Validity: 🥑	weight is unity.	r the spectrum should be scaled so that the maximum	
Preview	Parameters		PARAMETER	DESCRIPTION	
Black-Body Spectru	TemperatureKelvin	5700		The temperature in Kelvin of the black-body whose	_
Version: 1.0 Last Modified: Tuesday, Augus	☑ Normalize		TemperatureKelvin	radiation is simulated by the spectrum generated wit this snippet.	h
Black-body radiation is one of t			Normalize	This variable gives the user the option to scale the curso that the maximum weight is unity.	Irve
	Numerical Settings Size of Wavelength Window 1 µm Number Sampling Distance 5 nm	er Data Points 200 €			Close .::

# **Programmable Spectrum: Writing the Code**



# **Programmable Spectrum: Using Your Snippet**

Bear in mind that the function we have programmed only works for wavelength specification!	Edit Programmable Spectrum Generator Generator Settings	× <ul> <li>Specification In Frequency Domain</li> <li>Validity: ♥</li> </ul>	Modify your snippet again by clicking on Edit
Modify the sampling parameters according to the requirements of your simulation.	Parameters TemperatureKelvin ☑ Normalize	5700	You can modify the value of the global parameters you defined here         Snippet Help         Black-Body Spectrum         Version: 1.0         Last Modified: Tuesday, August 14, 2018         Black-body radiation is one of the best-known theories in physics, and it is employed across several fields. This snippet generates a spectrum according to the black body yshall radiate at, and whether the spectrum should be scaled so that the maximum weight is unity.
	Numerical Settings Size of Wavelength Window 1 µm Sampling Distance 5 nm	Number Data Points 200 -	PARAMETER       DESCRIPTION         TemperatureKelvin       The temperature, in Kelvin, of the black-body whose this simulated by the spectrum generated with this snippet.         Normalize       This variable gives the user the option to scale the curve so that the maximum weight is unity.

## **Programmable Spectrum: Output**

Field Vector Component At One Point	
Diagram Table Value at x-Coordinate	
Wavelength [µm]	0.9 1
	Field Vector Component At One Point

#### **Test the Code!**

```
Main Function
Complex value = new Complex(0, 0);
// Constants not included in Globals.
const double BoltzmannConstant = 1.3806505e-23;
const double ProportionalityConstantWienLaw = 2.8977729e-3;
if (Normalize) // Code to run if the curve is to be normalized.
{
   // Eq. (2) computes the wavelength at which the curve presents its maximum.
    double wavelengthMaximum = ProportionalityConstantWienLaw / TemperatureKelvin;
   // The normalization constant is equal to the value of the curve at wavelengthMaximum.
    double normalizationConstant = (Math.Pow(wavelengthMaximum, 5) * (Math.Exp((Globals.PlanckConstant *
        Globals.VacuumSpeedOfLight) / (wavelengthMaximum * BoltzmannConstant * TemperatureKelvin)) - 1)) /
        (8 * Math.PI * Globals.PlanckConstant * Globals.VacuumSpeedOfLight);
   // Eq. (1) multiplied by normalization constant gives the final value of S per wavelength.
    value = normalizationConstant * (8 * Math.PI * Globals.PlanckConstant * Globals.VacuumSpeedOfLight) /
        (Math.Pow(Position, 5) * (Math.Exp((Globals.PlanckConstant * Globals.VacuumSpeedOfLight) / (Position)
        * BoltzmannConstant * TemperatureKelvin)) - 1));
// Continued in next page.
```

#### **Test the Code!**

```
Main Function (continued)
// Continued from previous page.
else // Code to run if curve is not to be normalized.
{
   // Eq. (1) gives the value of S per wavelength.
    value = (8 * Math.PI * Globals.PlanckConstant * Globals.VacuumSpeedOfLight) /
        (Math.Pow(Position, 5) * (Math.Exp((Globals.PlanckConstant * Globals.VacuumSpeedOfLight) /
        (Position * BoltzmannConstant * TemperatureKelvin)) - 1));
}
// Eq. (1) is in dimensions of energy, and the programmable spectrum in VirtualLab must return field
// amplitudes:
value = Complex.Sqrt(value);
return value;
// End of code.
```

#### How to Use Your Custom Spectrum in a Source



#### **Document Information**

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VL version used for simulations	7.4.0.49
category	Feature Use Case
further reading	<ul> <li>How to Work with the Programmable Light Source And Example (Gaussian Beam)</li> <li>Programming a Chirped Gaussian Pulse Spectrum</li> </ul>